

Mathematics makes the unseen visible (and vice versa)

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Finnish Centre of Excellence
in Inverse Problems Research



What are inverse problems?

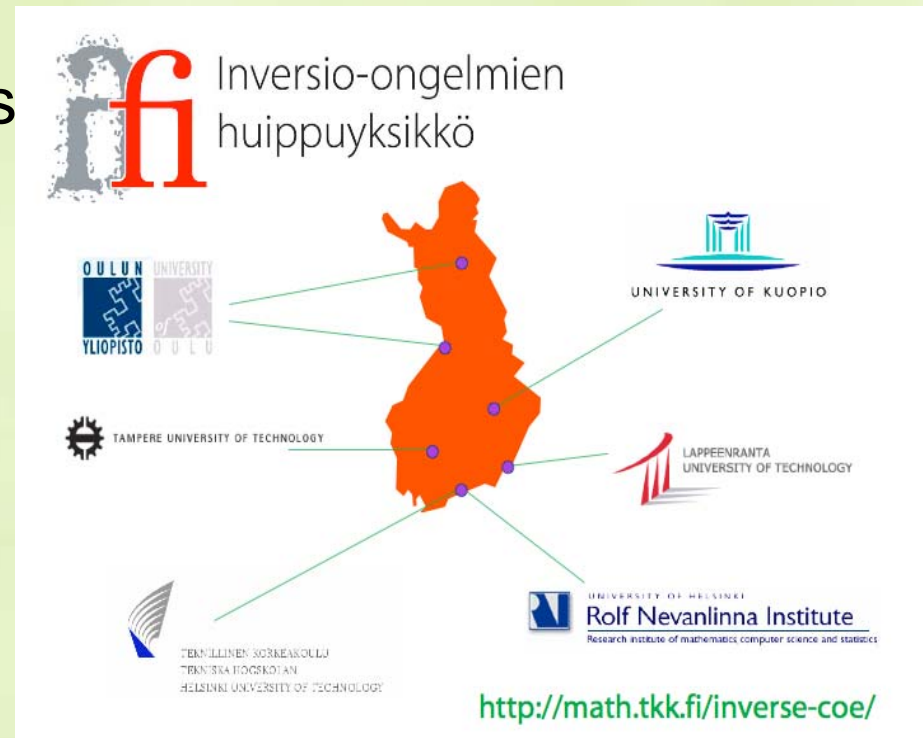
- ✿ From consequence to cause: what has caused the observations?
- ✿ The observations are explained with a mathematical model
- ✿ The data are usually indirect, limited and noisy
- ✿ Is the model unique and stable? How much can it vary?
- ✿ Much more difficult than the direct problem: mathematical analysis and the search for practical solution methods are essential
- ✿ Inverse problems = quest for information

Why is inverse difficult?

- ✿ Cf. computing in one's head: compute $15129^{1/2}$. The inverse function is more difficult than the direct one: $123^2 = (120+3)^2 = 15129$.
- ✿ Nonuniqueness of solution: Assume the cover of a well and its (convex) aperture in the plane (2D). Which cover shapes never slip through the planar aperture in any orientation (3D)? *Hadamard 2: Uniqueness of the inverse mapping?*
- ✿ Errors: Even though the formal inverse solution exists, the smallest error in data can lead to an unbounded error in the solution. *Hadamard 3: Continuity of the inverse mapping?*
- ✿ How do we know what is the maximal information derivable from the data, and what is the corresponding mathematical model?

IP-Centre of excellence

- ❖ 6 universities, everything from viruses to galaxies, from pure mathematics to industrial applications
- ❖ Wide international network
- ❖ World leader
- ❖ Long experience in mathematical methods in various applications

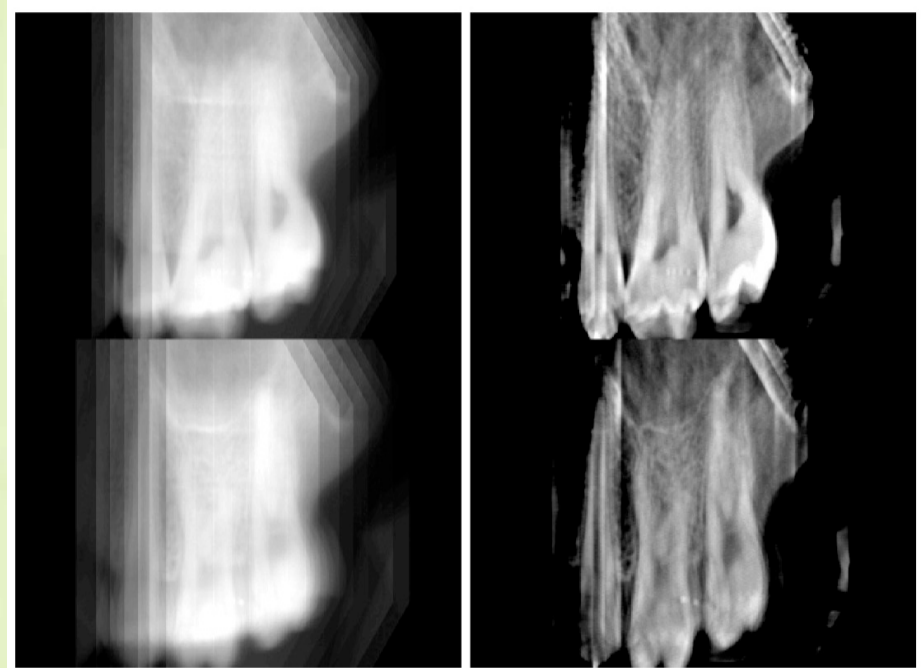


Research areas at TUT

- ✿ Space research, dynamical systems, climate/ecological studies (carbon cycle+biomass), remote sensing (laser scanning), biomedical imaging
- ✿ Generalized projection operators, phase-space tomography, Poincare inverse problem of dynamics, optimal combination of multiple data modes (maximum compatibility estimate, MCE), information content of data, model selection/reduction/expansion, efficient algorithms for large datasets

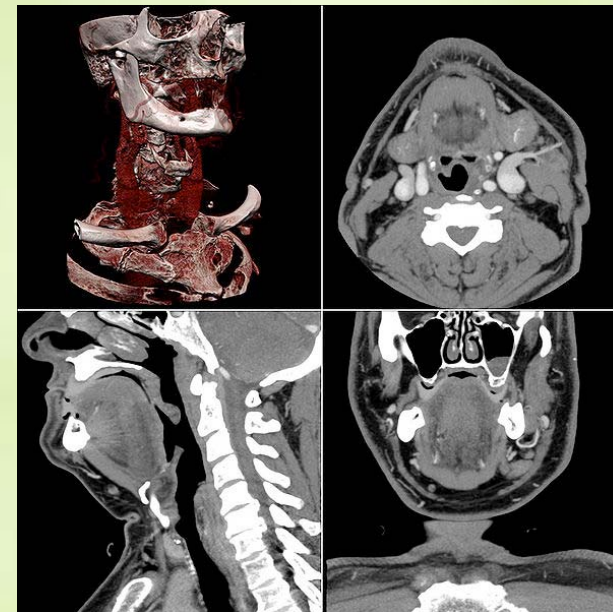
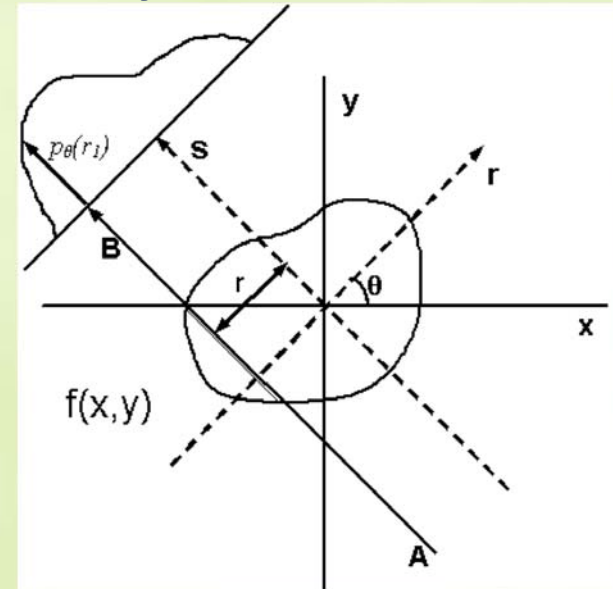
Dental tomography

- ✿ Limited angle X-ray data
- ✿ Complete 3-D model of tooth
- ✿ Statistical inversion yields a better solution than the traditional X-ray projection (Bayes: *a posteriori* - distribution includes *a priori* -information)



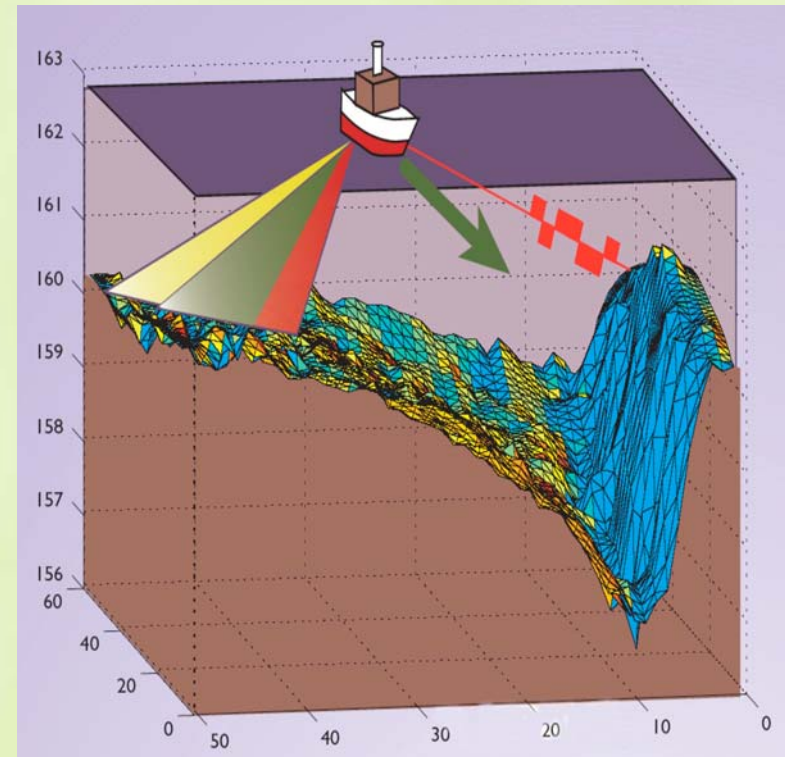
Biomedical tomography

- ✿ Radon transform ~100 y ago started the mathematics of imaging (X-ray data)
- ✿ Many tomographic methods: MRI, PET, etc.
- ✿ New: EIT, diffusion tomography
- ✿ EIT is based of the mathematical solution of the *Calderon problem* (2000's): electric conductivity inside the the target is determined from voltage measurements on the boundary
- ✿ Breast cancer, pulmonary activity: tissues conduct electricity in different ways



Sonar and radar

- ✿ Full 3-D sonar (side directions included)
- ✿ Complete topographic model of river bottom
- ✿ Pulse coding and inversion methods essential (can even derive *perfect codes*)
- ✿ Mathematics makes weather radar work even in difficult conditions



Space radars



- ✿ Sodankylä EISCAT 32 m
- ✿ Arecibo 300 m
- ✿ Ionosphere, space debris, moon, asteroids

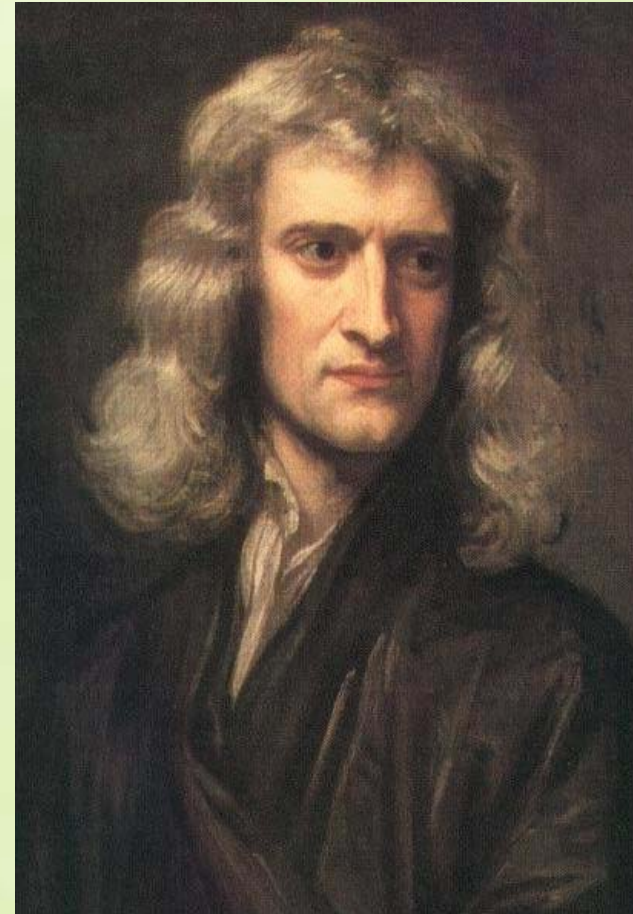
Galaxies and dark matter

- ✿ We live at the edge of the disk
- ✿ Dust obscures a lot
- ✿ But high density of material: lots of data
- ✿ How is *dark matter* distributed?
What is it made of?
- ✿ Thus: what potential field (\Rightarrow distribution of matter) holds the system together?



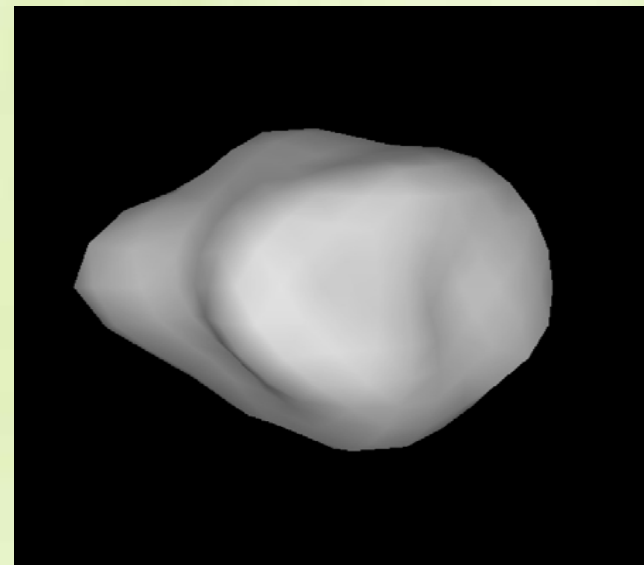
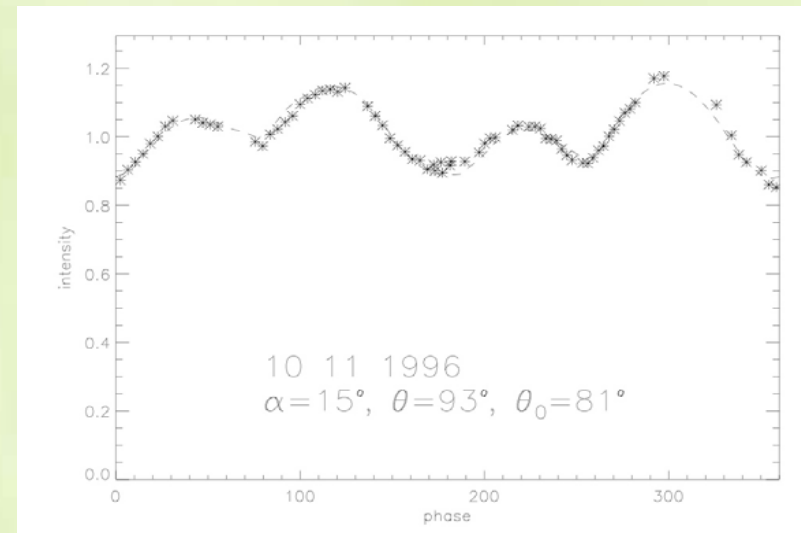
Generalization of Newton's inverse problem

- ✿ $F \propto 1/r^2 \Rightarrow$ elliptic orbit
(conical sections)
- ✿ Inverse problem:
elliptic orbit $\Rightarrow 1/r^2$
(when motion around foci)
- ✿ How to generalize this
to $N \gg 2$ bodies (e.g.
positions and velocities
of a billion stars)?
- ✿ Tomography in six
dimensions!
- ✿ Poincare IP of
dynamics: torus
construction



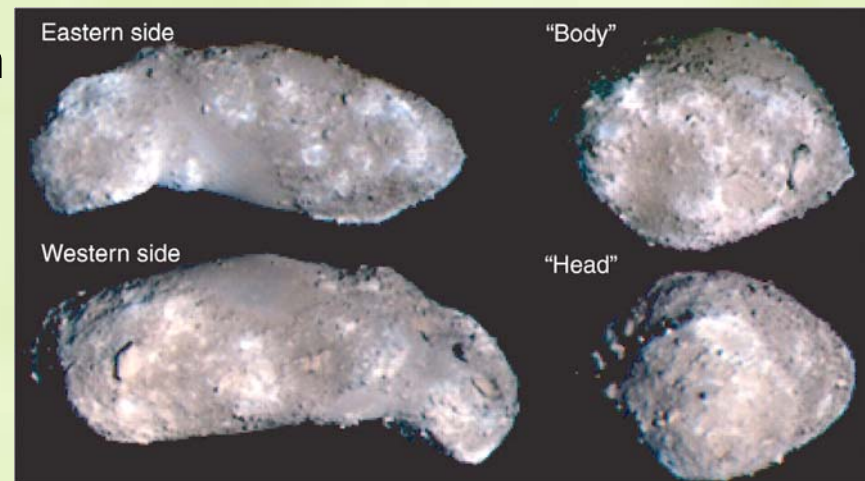
Space research

- ✿ Brightness measurements (lightcurves) of asteroids at various times (as the target moves in its orbit and rotates)
- ✿ What kind of an object created the data?
- ✿ Complete model of shape, surface and spin

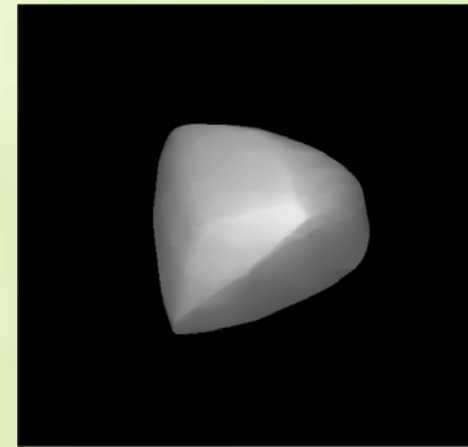
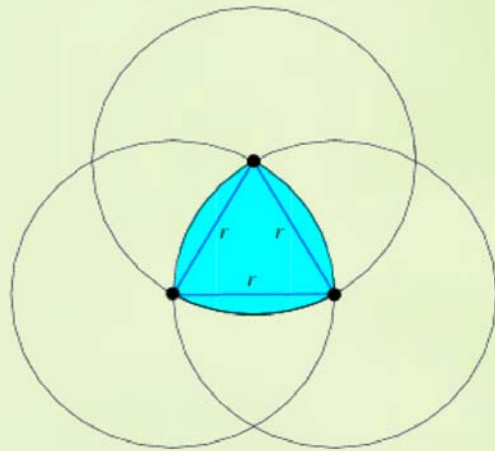


Solar system full of asteroids

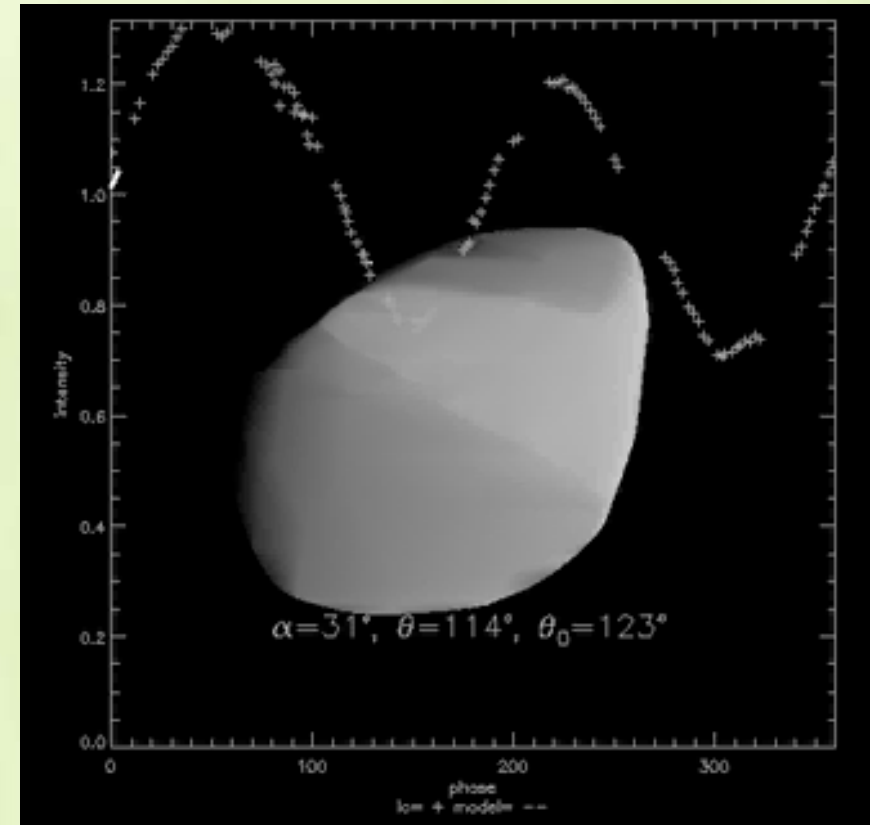
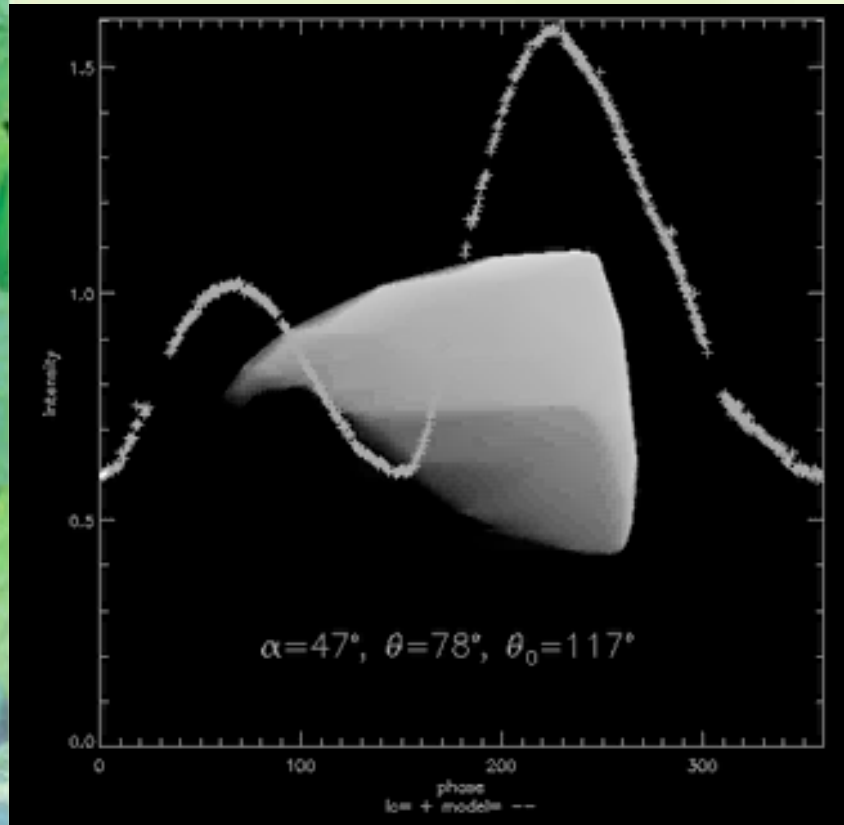
- ✿ Millions of objects ~1-100 km-class, innumerable smaller ones
- ✿ Each a world of its own
- ✿ “Jurassic Park”: asteroids tell us about the birth and evolution of the solar system
- ✿ Similar populations around other stars
- ✿ Direct images only of a few asteroids



Mathematics maps asteroids



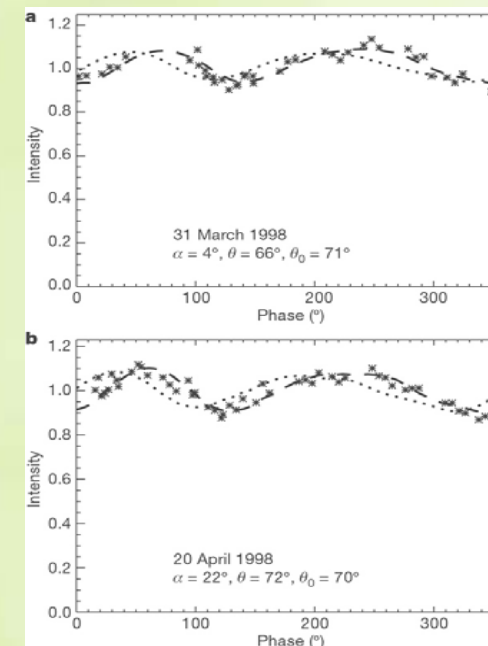
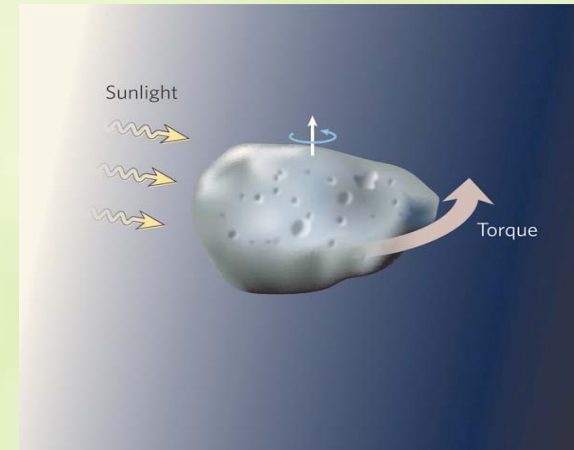
- ✿ *Generalized projections*: inverse problem
- ✿ Usual areas of projections insufficient (cf. Reuleaux triangle)
- ✿ Various viewing and illumination directions yield a both unique and stable solution with convexity constraint
- ✿ Tested in practice in several ways
- ✿ Was considered unsolvable...



- ✿ Reconstruction possible with various observing directions
- ✿ Inexpensive data: often even a small telescope and CCD camera (+software) sufficient

New discoveries: Solar motor

- ✿ *YORP effect*: solar-accelerated rotation was found from the lightcurves of 1862 Apollo in 2006 [Kaasalainen et al., *Nature* 446, 420 (2007)]
- ✿ Several other YORP targets found in the same way after Apollo
- ✿ Explains binary asteroids
- ✿ “Rubble-pile” asteroid breakup, change of shape; change of orbit (*Yarkovsky-effect*)



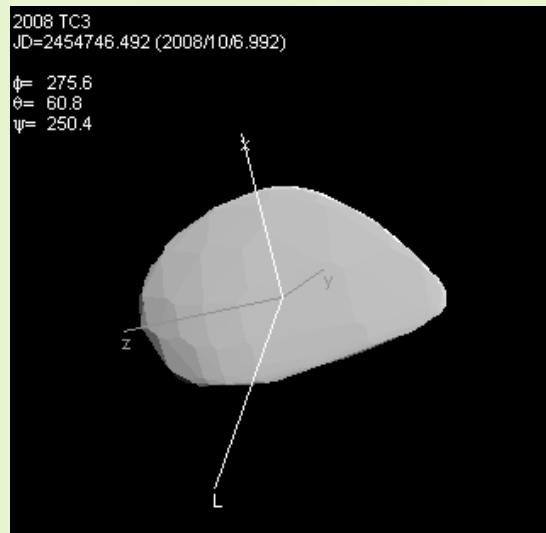
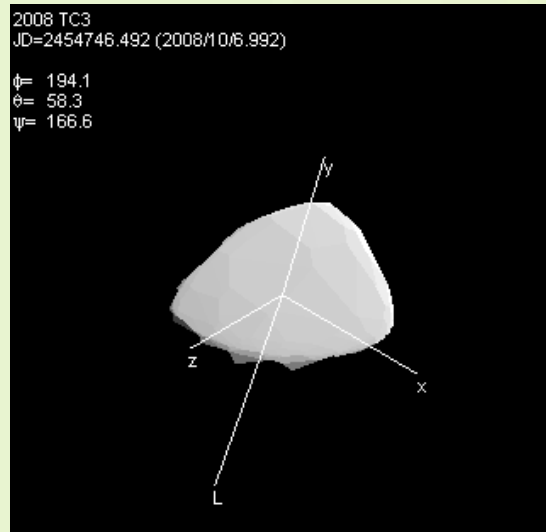
“New” phenomenon in solar system

☀ Sunlight affects the motion of asteroids, including targets delivered into Earth-crossing orbits

☀ Can even be used for avoiding collisions (let's paint the asteroid suitably!)



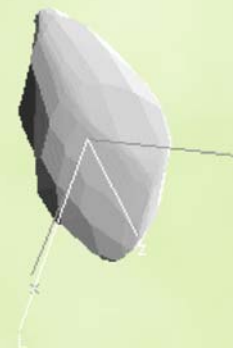
One that fell to Earth: 2008TC3



- ✳ Small (less than 10 m, mass 100 tons), disintegrated in atmosphere
- ✳ Pieces found in Africa, so we know the composition
- ✳ And we know the orbit; first predicted collision with Earth...
- ✳ Brightness data were obtained just before collision: shape and rotation (precession)

2008 TC3
JD=2454746.492 (2008/10/6.992)

$\phi = 131.7$
 $\theta = 60.9$
 $\psi = 105.2$

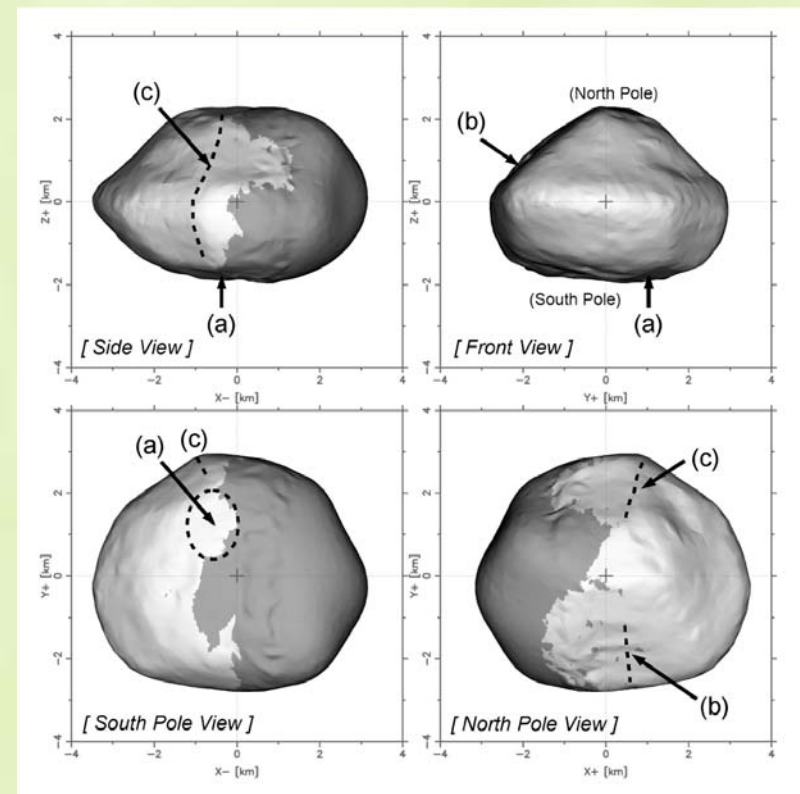


We map the solar system

- ✿ N*10 000 mapped asteroids within next 10 years
- ✿ The only way of removing the last big “white spot” of the solar system
- ✿ Databases and data acquisition international and Internet-based
- ✿ Several survey-progs >2010: Pan-STARRS, LSST, Gaia, etc.
- ✿ Shape@home screen saver: map your own asteroid!

Space research: how to see the dark side

- ✿ ESA's Rosetta probe flew fast by the asteroid Steins and saw only one side (and the hi-res camera was lost)
- ✿ We reconstructed the other half from brightness data
- ✿ Stable solution: strong constraints by images
- ✿ YORP effect visible



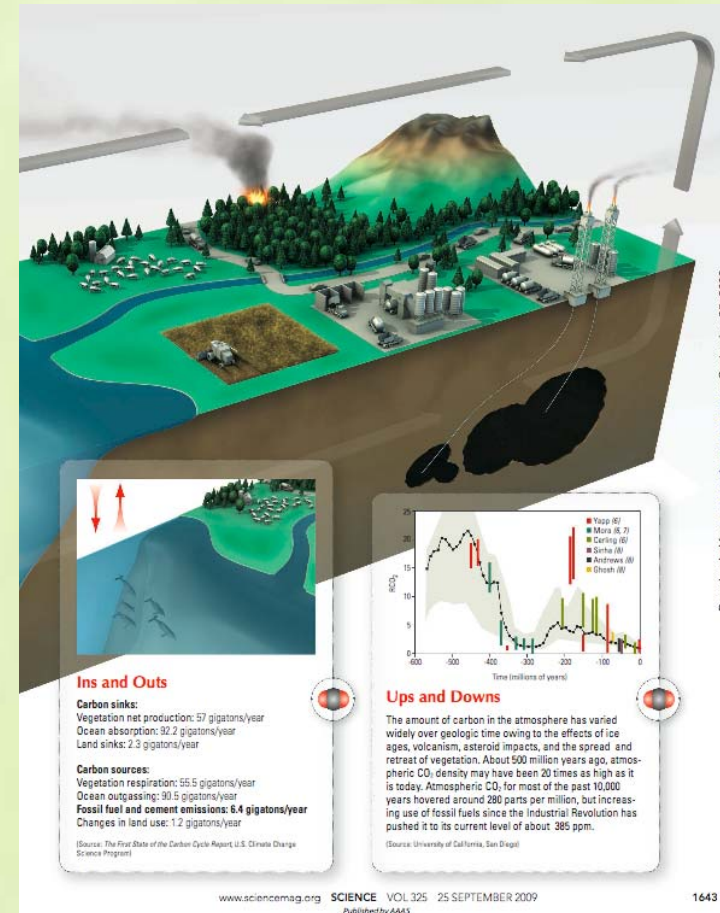
Rosetta: Steins flyby

- ✿ Flyby lasted only 7 minutes
- ✿ Closest encounter 800 km, target size 5 km
- ✿ Speed w.r.t. target 8.6 km/s = 31000 km/h
- ✿ Dark side reconstruction cost ~0 €, gave significant added value to a project of 700 million €
- ✿ Keller et al., Science 2010 (Osiris team, ~50 co-authors)



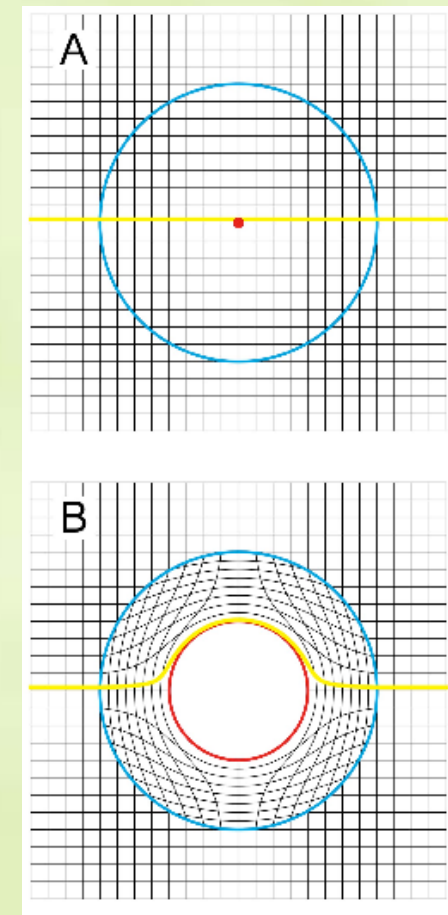
Carbon cycle, biomass estimation+carbon footprint

- ✿ Complex system, model prone to errors
- ✿ Model selection problem, design of experiments
- ✿ Biomass estimation, disintegration profiles
- ✿ Laser scanning technologies, mobile platforms?



Invisibility cloaking

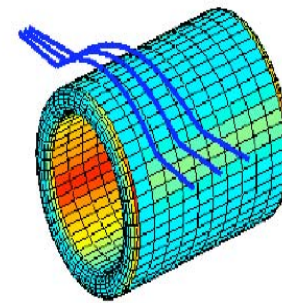
- ✿ Light circumvents the target: looks as if nothing obstructed its path
- ✿ Inverse problem: what kind of a model of the medium totally removes observations of the target?



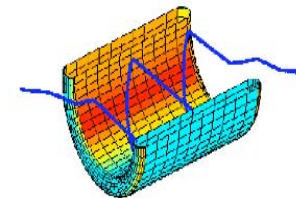
Optical wormhole

- ✿ Light travels in an invisible tube (“Harry Potter’s sleeve”): only the exit is seen
- ✿ A way to make, say, a 3-dimensional television
- ✿ Metamaterials (negative refraction coefficient) in the tube/cloak

Ray tracing simulations:



Rays travelling outside.



A ray travelling inside.